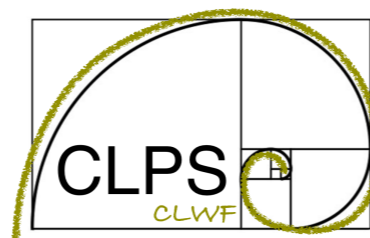


# Albert en het Universum

Gustaaf C. Cornelis



Vrije  
Universiteit  
Brussel



Universiteit  
Antwerpen

# Onderwerpen

- bewijzen voor de ART op kosmologische schaal
- implicaties van de ART op kosmologische schaal
- de geciteerde Einstein

Even recapitulieren

- $G^{\mu\nu} \sim T^{\mu\nu}$
- $G^{\mu\nu} = \kappa T^{\mu\nu}$ 
  - $\kappa = 8\pi G/c^4$ 
    - $\kappa$ : constante van Einstein (kappa)
    - $G$ : constante van Cavendish:  $6,6754 \times 10^{-11} \text{ m}^3 \text{ s}^{-2} \text{ kg}^{-1}$
  - $G^{\mu\nu}$ : Einstein-tensor
    - uitdrukking voor kromming (geometrie)
  - $T^{\mu\nu}$ : Energie/impuls-tensor
    - uitdrukking voor massaverdeling, oorzaak van gravitatie



Space acts on matter, telling it *how to move*.  
In turn, matter reacts back on space, telling it  
*how to curve*.

Misner, C.W., Thorne, K.S. & Wheeler, J.A., 1973, *Gravitation*. San Francisco, W.H. Freeman, p.5

Empirische bewijzen

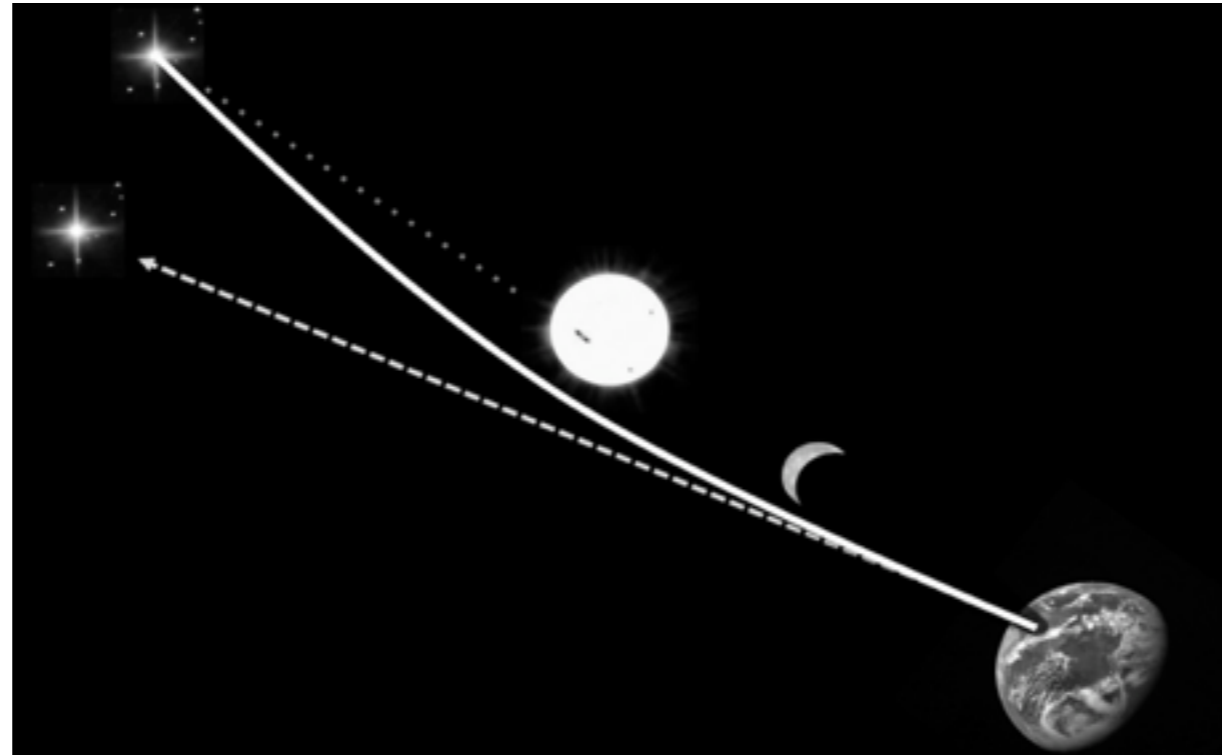


<http://www.elimparcial.es/noticia/23558/sociedad/>

# Eddingtons vermeende bewijs (1917)



<http://link.springer.com>



<http://astrobob.areavoices.com>



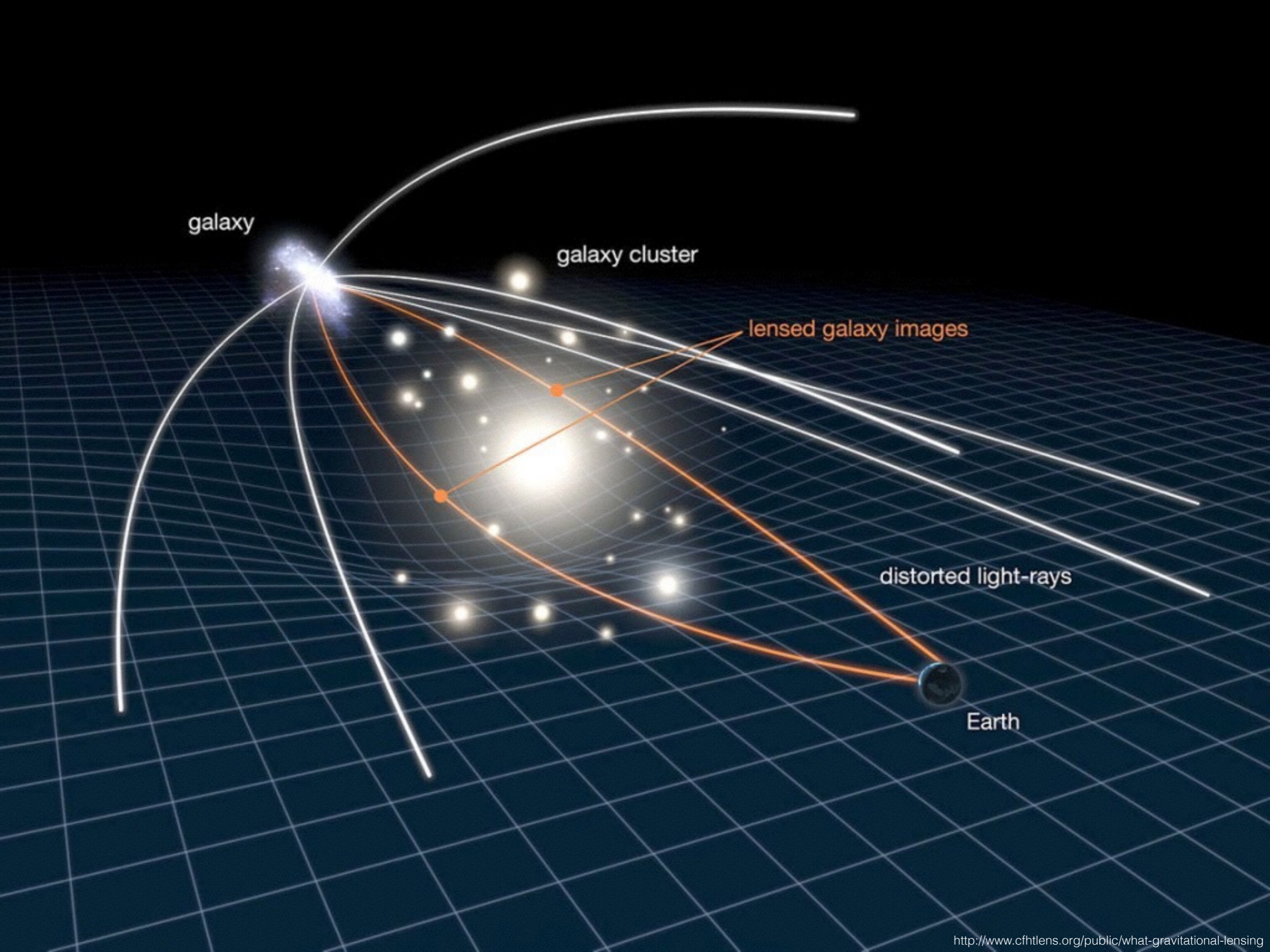
Of course, there is no hope of observing this phenomenon directly.



Einstein, A., 1936, "Lens-like action of a star by the deviation of light in the gravitational field". *Science* 84, pp.506-507.

# Tweeling Quasar (1979)





galaxy

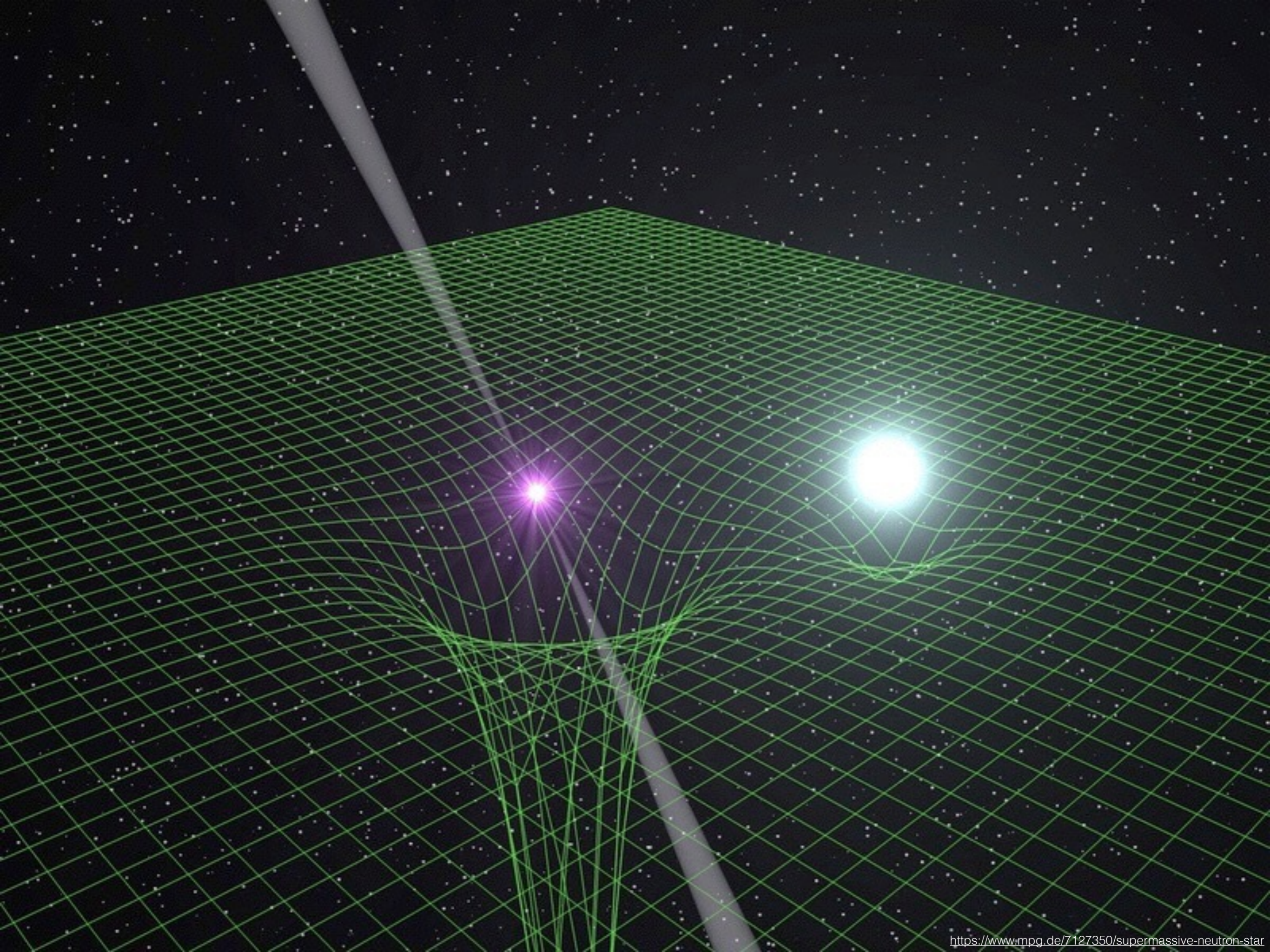
galaxy cluster

lensed galaxy images

distorted light-rays

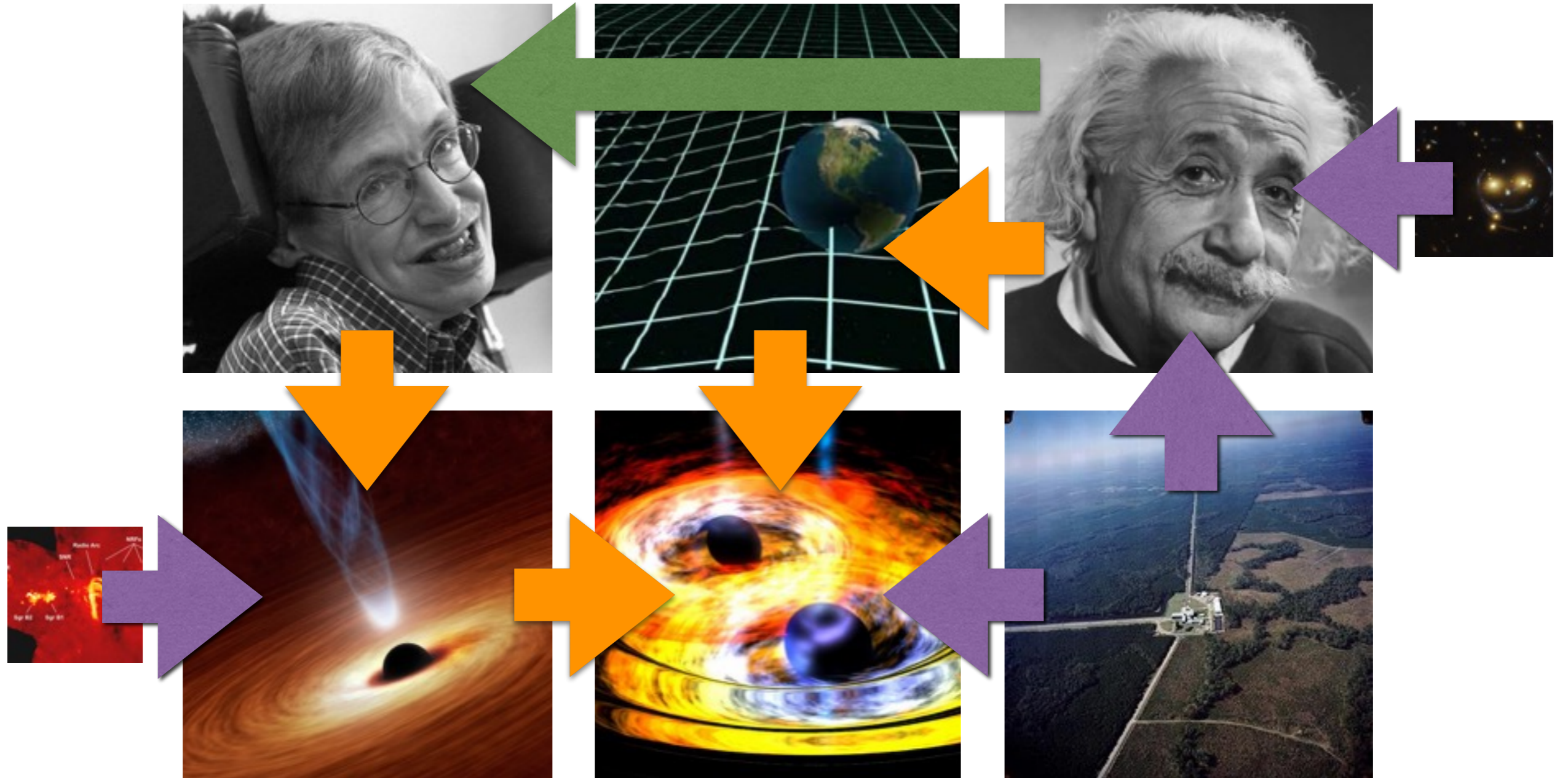
Earth

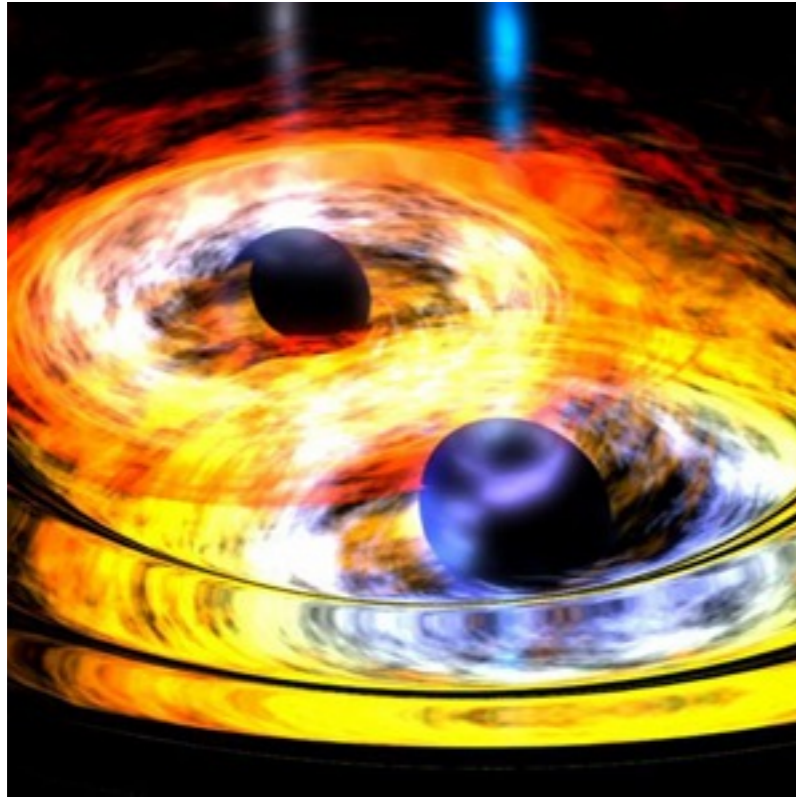




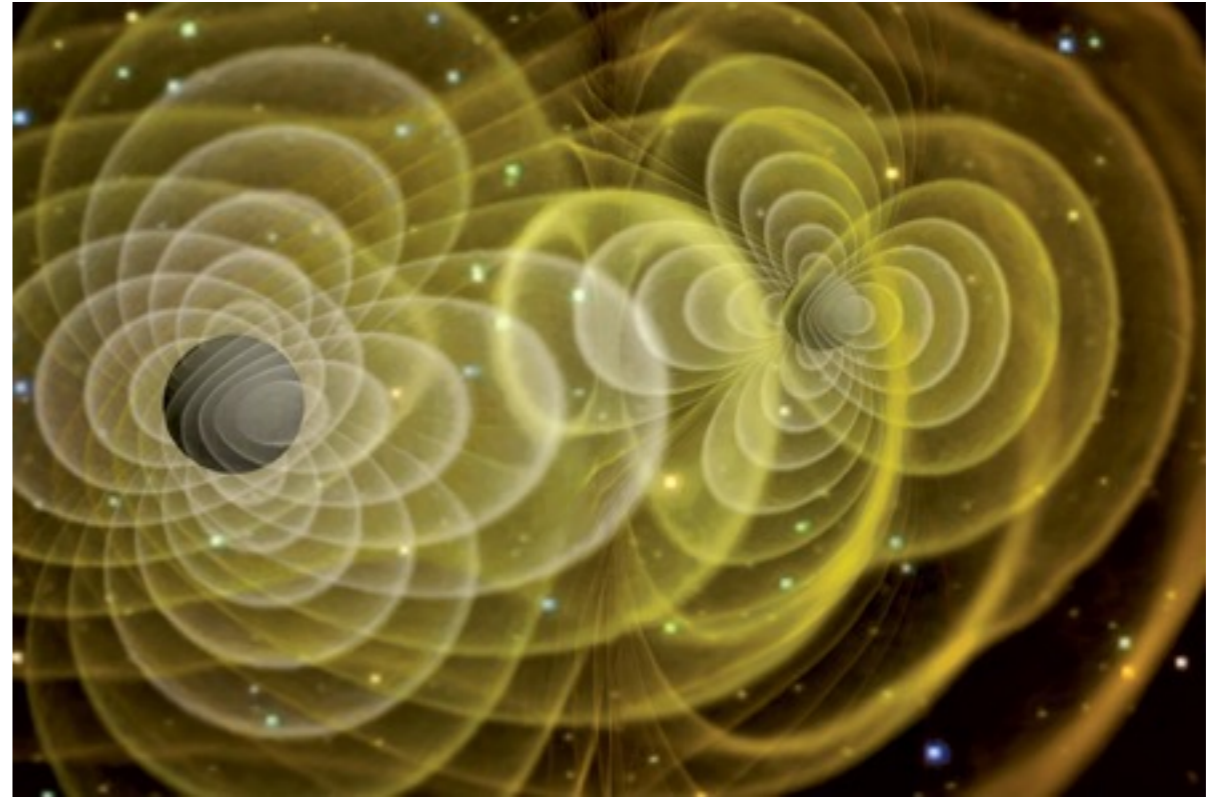


# Gravitationegolven

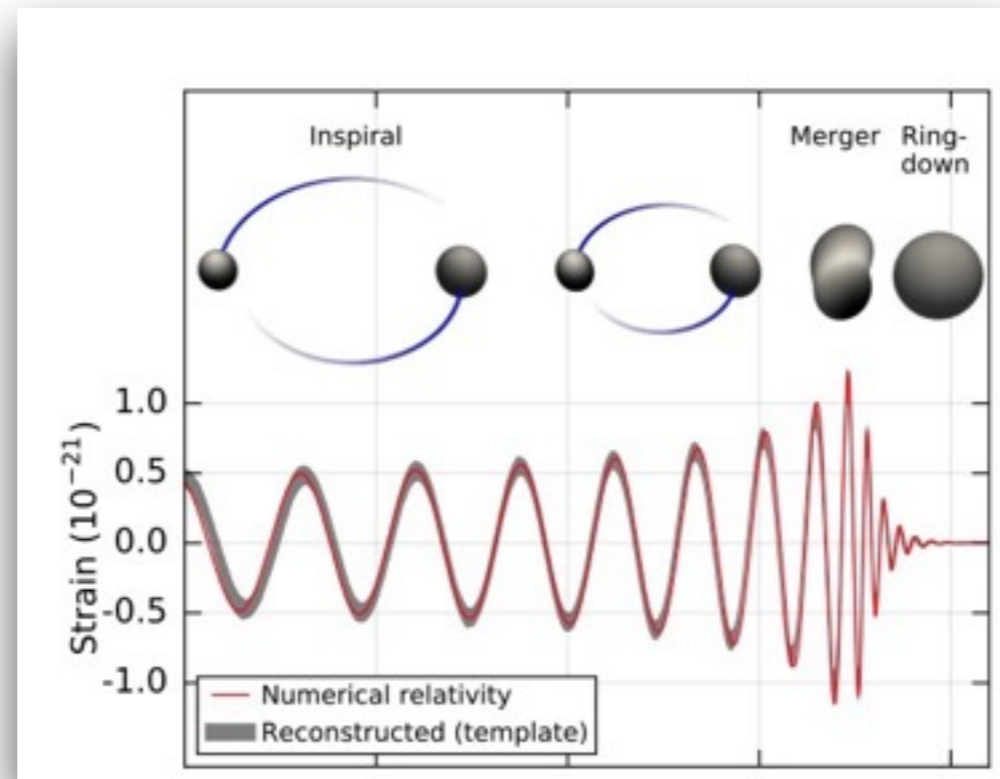




<http://physicsworld.com/cws/article/news/2016/feb/11>

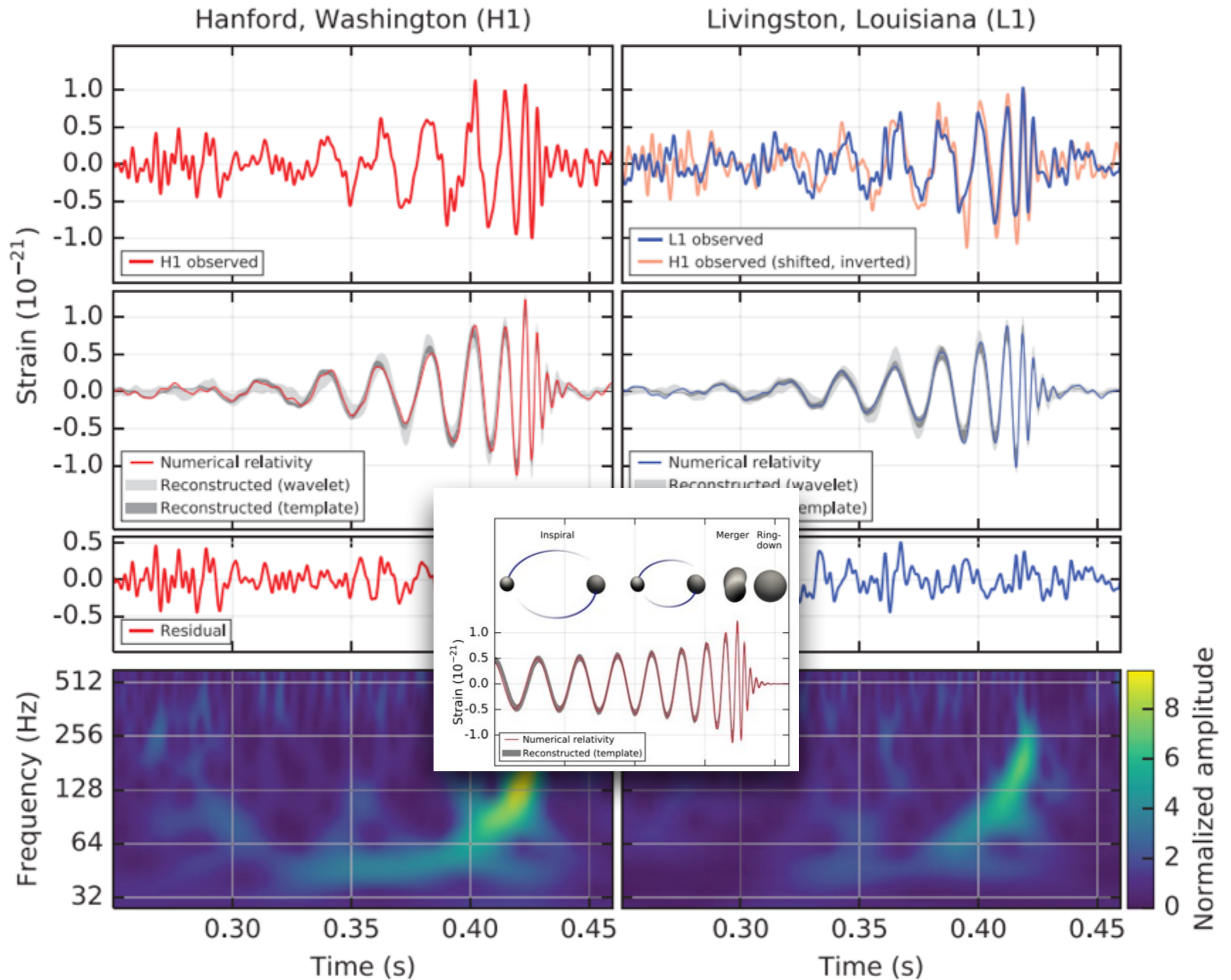


<https://www.newscientist.com/article/2076754>



<http://journals.aps.org/prl/pdf/10.1103/PhysRevLett.116.061102>





# Relativistische Kosmologie

# Artikel uit 1917

## Kosmologische Betrachtungen zur allgemeinen Relativitätstheorie.

VON A. EINSTEIN.

Es ist wohlbekannt, daß die Poisson'sche Differentialgleichung

$$\Delta \phi = 4\pi K \rho \quad (1)$$

in Verbindung mit der Bewegungsgleichung des materiellen Punktes die Newton'sche Fernwirkungstheorie noch nicht vollständig ersetzt. Es muß noch die Bedingung hinzutreten, daß im räumlich Unendlichen das Potential  $\phi$  einem festen Grenzwerte zustrebt. Analog verhält es sich bei der Gravitationstheorie der allgemeinen Relativität; auch hier müssen zu den Differentialgleichungen Grenzbedingungen hinzutreten für das räumlich Unendliche, falls man die Welt wirklich als räumlich unendlich ausgedehnt anzusehen hat.

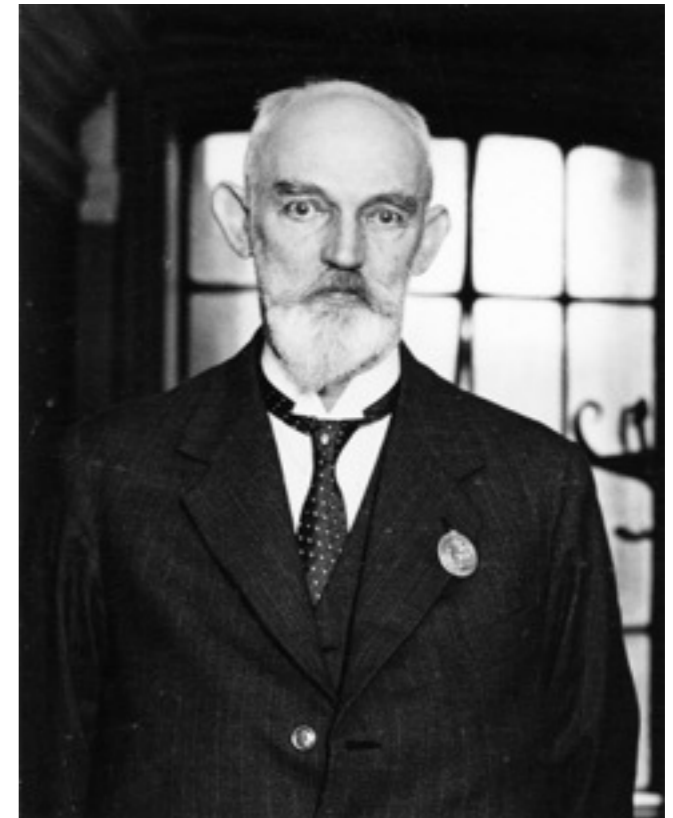
Bei der Behandlung des Planetenproblems habe ich diese Grenzbedingungen in Gestalt folgender Annahme gewählt: Es ist möglich, ein Bezugssystem so zu wählen, daß sämtliche Gravitationspotentiale  $g_{\alpha}$  im räumlich Unendlichen konstant werden. Es ist aber a priori durchaus nicht evident, daß man dieselben Grenzbedingungen ansetzen darf, wenn man größere Partien der Körperwelt ins Auge fassen will. Im folgenden sollen die Überlegungen angegeben werden, welche ich bisher über diese prinzipiell wichtige Frage angestellt habe.

### § 1. Die Newton'sche Theorie.

Es ist wohlbekannt, daß die Newton'sche Grenzbedingung des konstanten Limes für  $\phi$  im räumlich Unendlichen zu der Auffassung hinführt, daß die Dichte der Materie im Unendlichen zu null wird. Wir denken uns nämlich, es lasse sich ein Ort im Weltraum finden, um den herum das Gravitationsfeld der Materie, im großen betrachtet, Kugelsymmetrie besitzt (Mittelpunkt). Dann folgt aus der Poisson'schen Gleichung, daß die mittlere Dichte  $\rho$  rascher als  $\frac{1}{r^2}$  mit wachsender Entfernung  $r$  vom Mittelpunkt zu null herabsinken muß, damit  $\phi$  im

# Opdracht door Willem de Sitter

- Wat als we alle massa in het universum in rekening brengen?
- Hoe kromt het universum zichzelf?



<http://alchetron.com/Willem-de-Sitter-1231398-W>

- $G_{\mu\nu} = \kappa T_{\mu\nu}$

- laat enkel dynamische modellen voor het universum toe

- $G_{\mu\nu} + \lambda g_{\mu\nu} = \kappa T_{\mu\nu}$

- invoering  $\lambda$

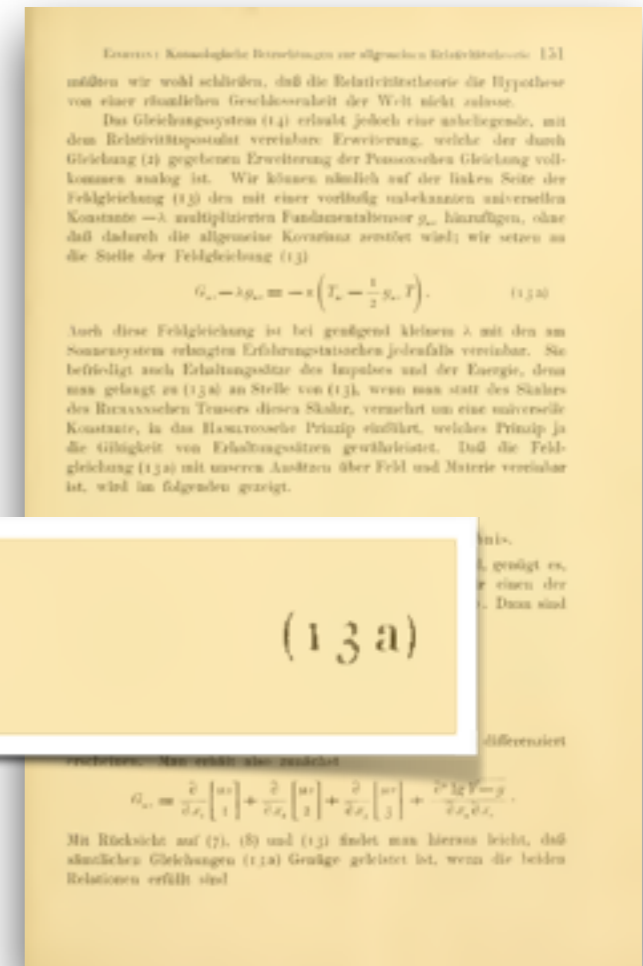
$$G_{\mu\nu} - \lambda g_{\mu\nu} = -\kappa \left( T_{\mu\nu} - \frac{1}{2} g_{\mu\nu} T \right). \quad (13a)$$

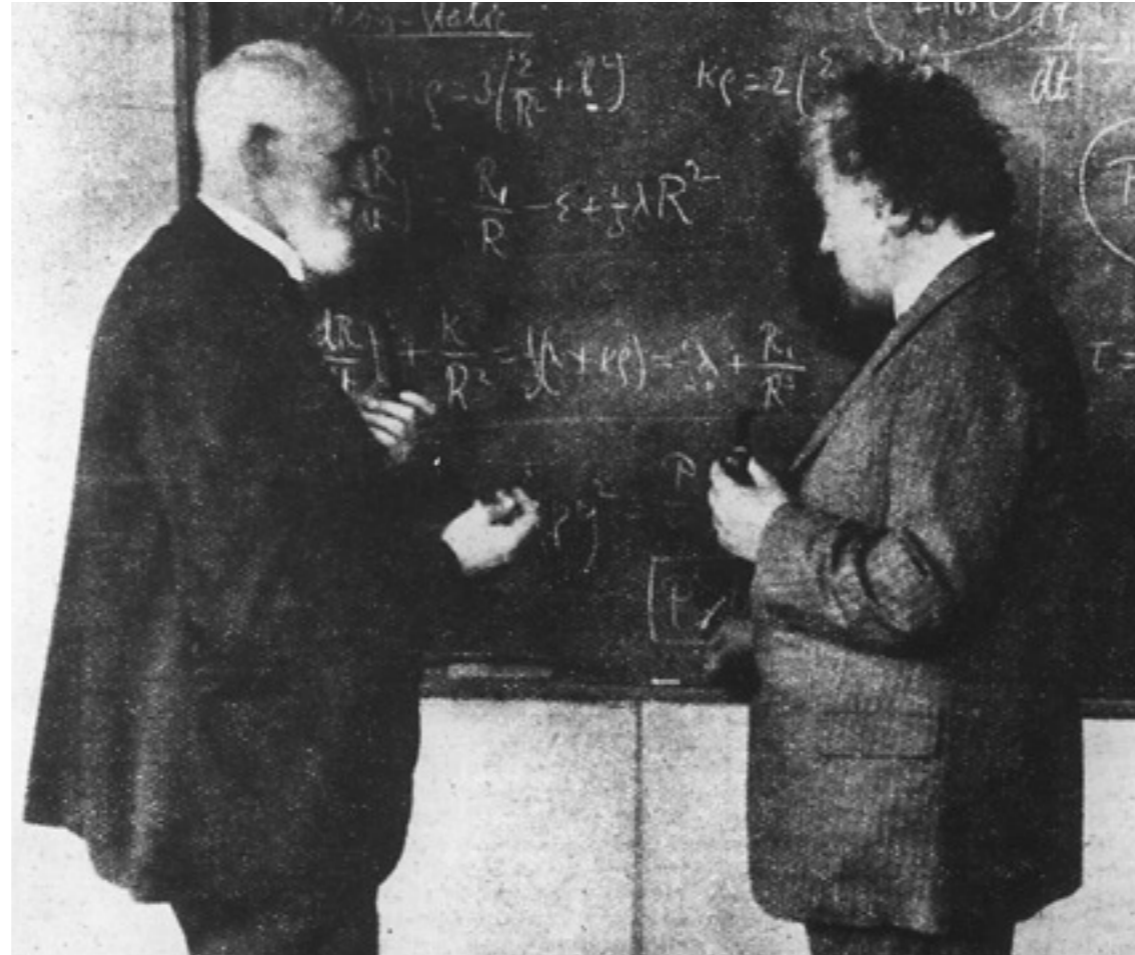
- kosmologische constante

- met geschikte waarde een statisch model

- kosmologisch principe

- universum is isotroop en homogeen





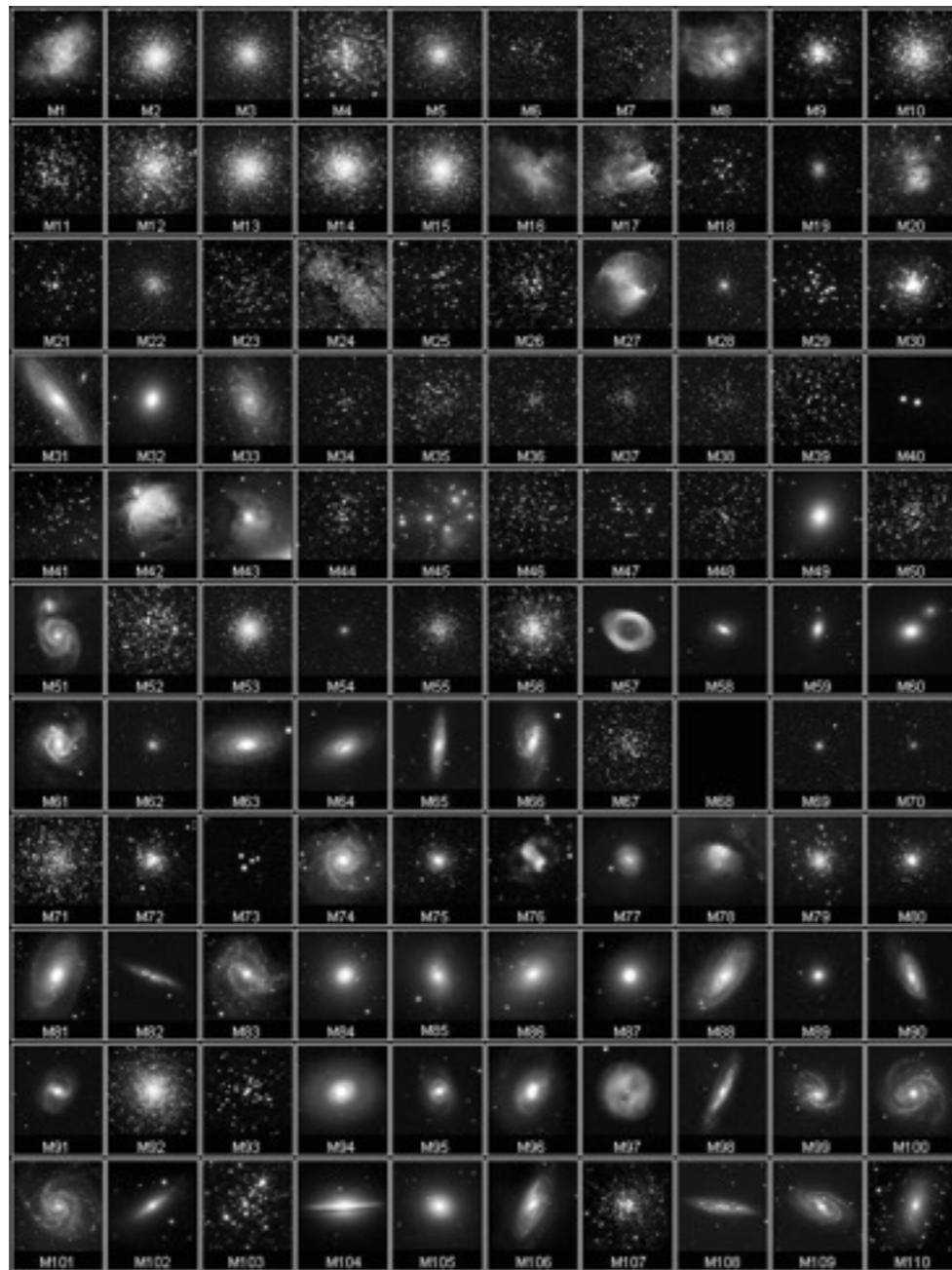
[http://forum.lametayel.co.il/forum\\_getfile\\_large/id-77745/](http://forum.lametayel.co.il/forum_getfile_large/id-77745/)

Ik wil benadrukken dat uit mijn resultaat een positieve kromming van het universum [implosie] volgt, zelfs als ik de extra term [met lambda] niet toevoeg. Die term is alleen maar nodig om een statisch model mogelijk te maken zoals de lage snelheden van de sterren vereisen.

Ster van Kapteyn:  
245 km/s

Die theoretische Auffassung der tatsächlichen Welt wäre also, falls dieselbe unserer Betrachtung entspricht, die folgende. Der Krümmungscharakter des Raumes ist nach Maßgabe der Verteilung der Materie zeitlich und örtlich variabel, läßt sich aber im großen durch einen sphärischen Raum approximieren. Jedenfalls ist diese Auffassung logisch widerspruchsfrei und vom Standpunkte der allgemeinen Relativitätstheorie die naheliegendste: ob sie, vom Standpunkt des heutigen astronomischen Wissens aus betrachtet, haltbar ist, soll hier nicht untersucht werden. Um zu dieser widerspruchsfreien Auffassung zu gelangen, mußten wir allerdings eine neue, durch unser tatsächliches Wissen von der Gravitation nicht gerechtfertigte Erweiterung der Feldgleichungen der Gravitation einführen. Es ist jedoch hervorzuheben, daß eine positive Krümmung des Raumes durch die in demselben befindliche Materie auch dann resultiert, wenn jenes Zusatzglied nicht eingeführt wird: das letztere haben wir nur nötig, um eine quasi-statische Verteilung der Materie zu ermöglichen, wie es der Tatsache der kleinen Sternengeschwindigkeiten entspricht.

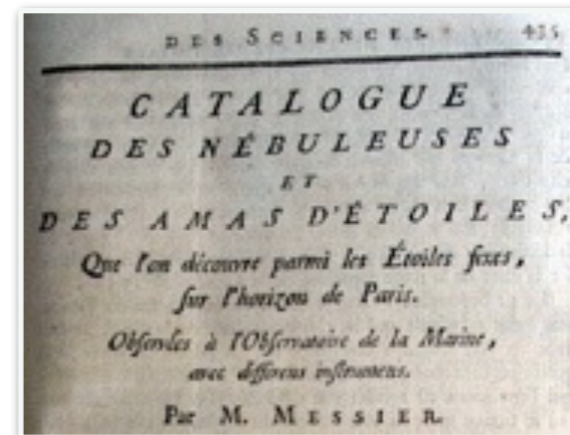
# Nebulae



<http://messier.seds.org/>

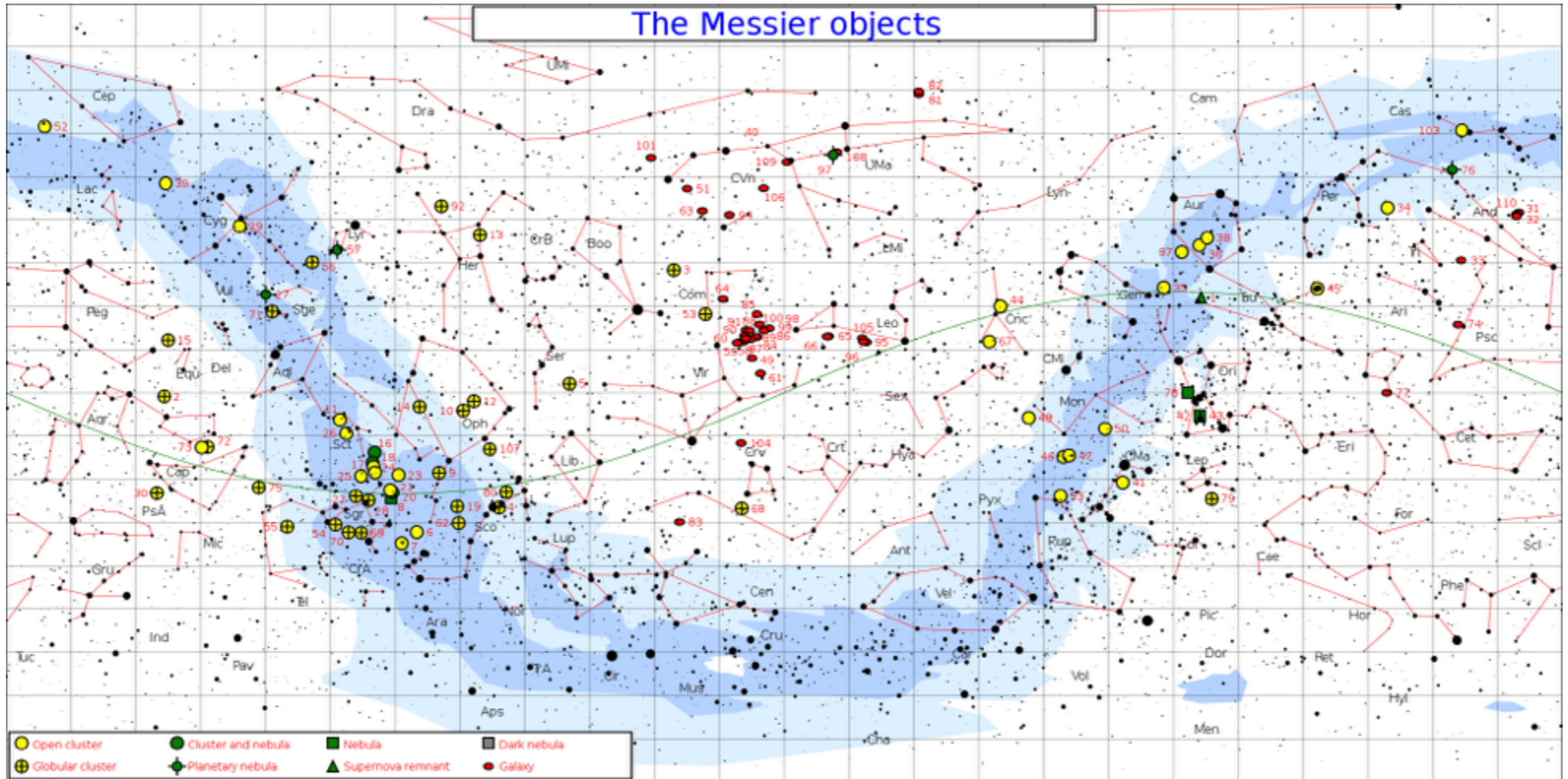
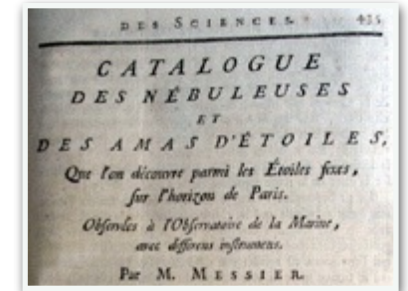


<http://www.nasa.gov/>





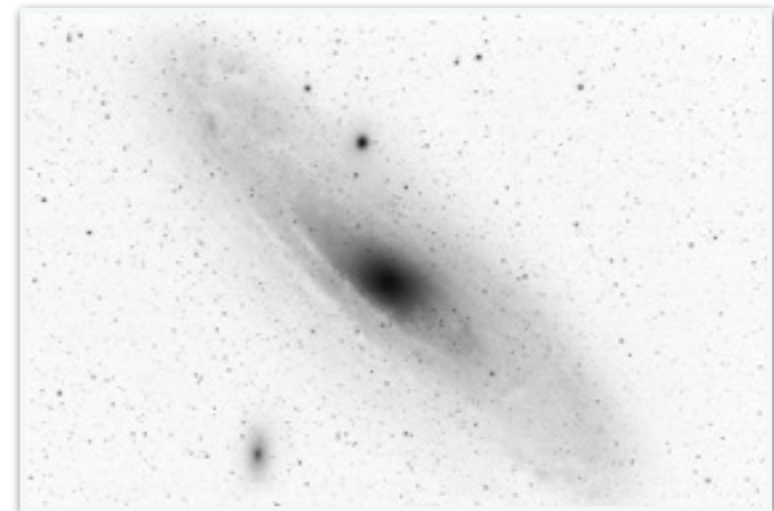
# 111 Messier Objects





<http://alchetron.com/Vesto-Slipher-1239034-W>

- Vesto Melvin Slipher
- 1912-1914
- nebulae roteren
- nebulae bewegen zich tegen hoge snelheid van ons weg
- met één uitzondering (M31)



<http://www.toledoastronomy.org/>

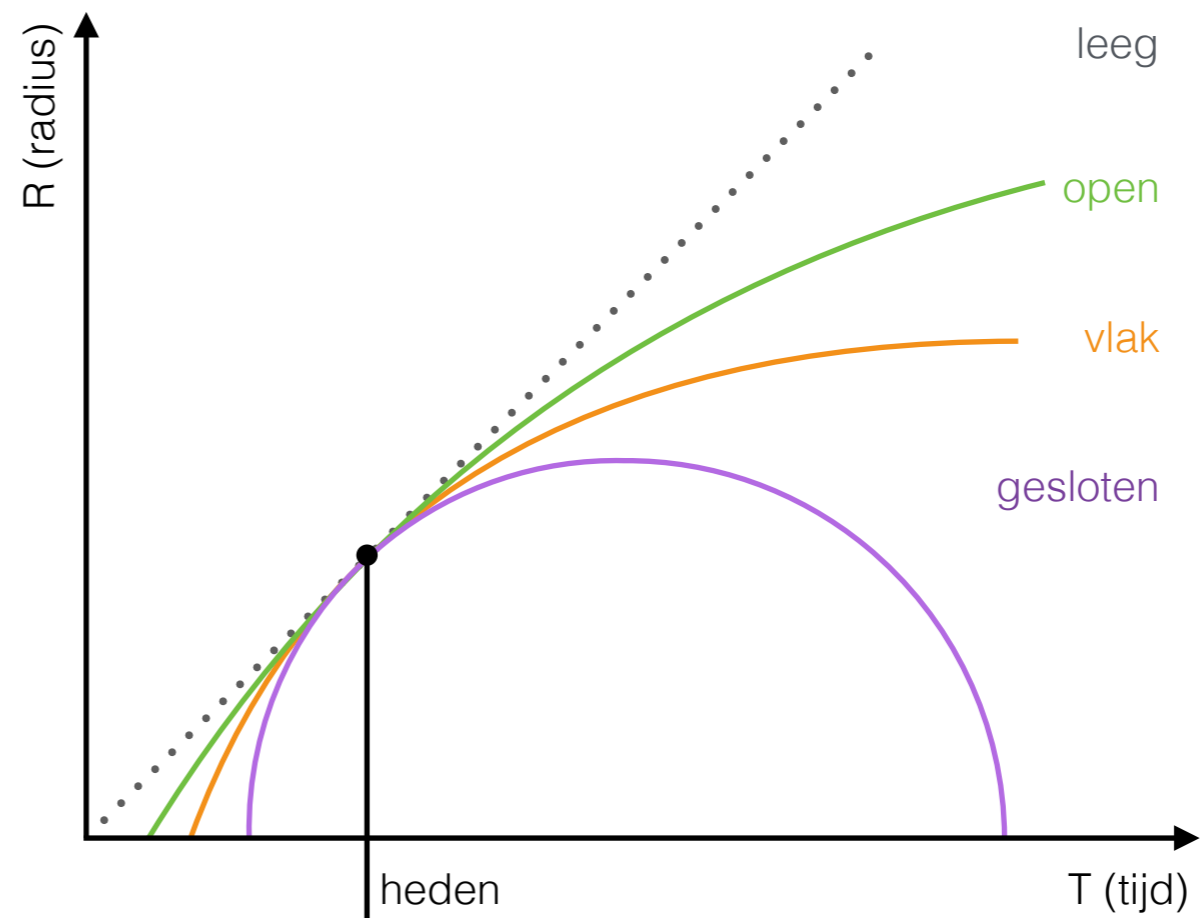
# Toestand 1917

- Geen idee van de afstanden tot nebulae
- Het universum is lokaal quasi-statisch
- Een heelalmodel betreft het ganse universum

# Friedmann modellen (1922-1924)



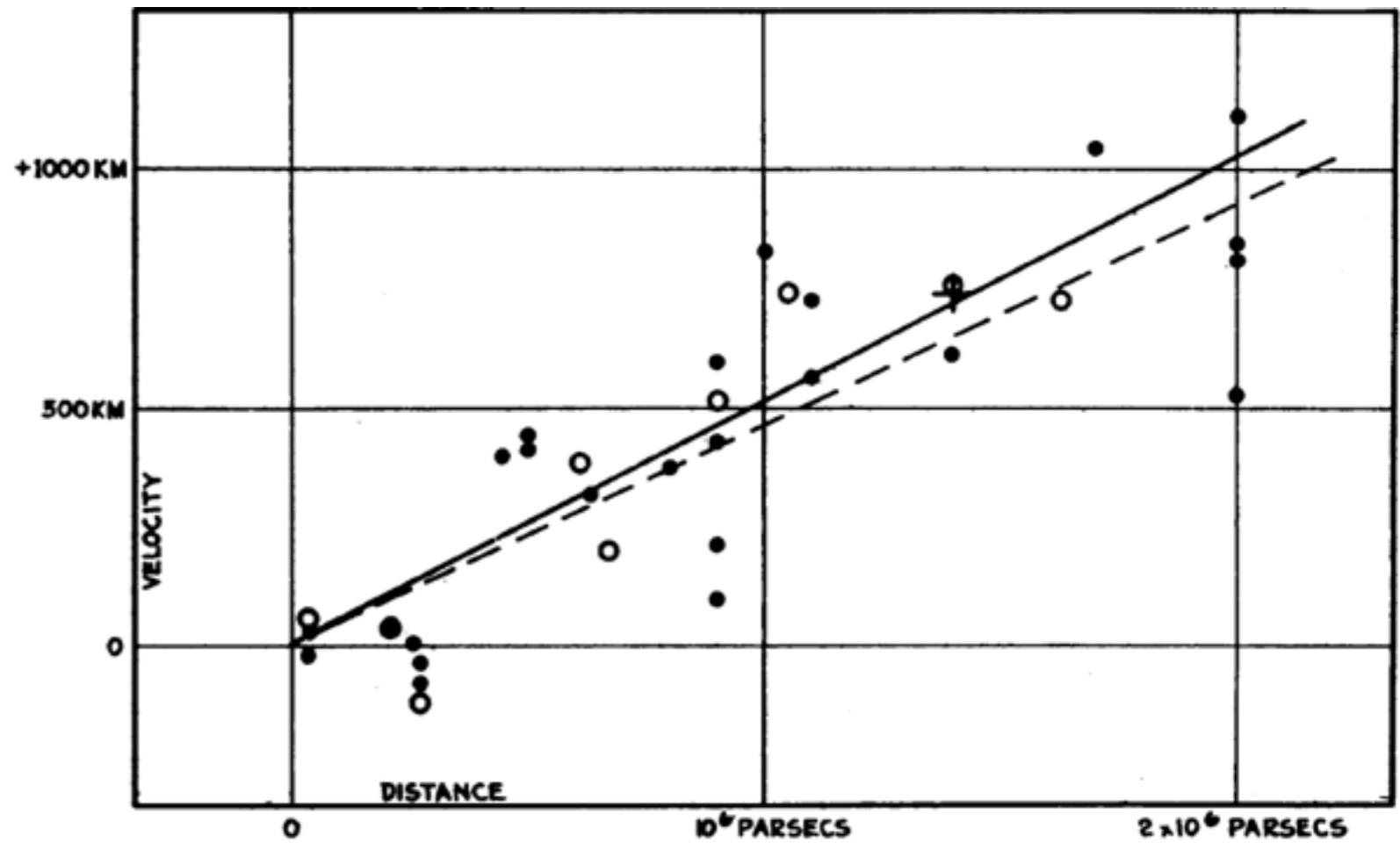
<http://www.frodo.at/Alexander-Friedmann.html>



# Hubbles ontdekkingen (1923-1929)

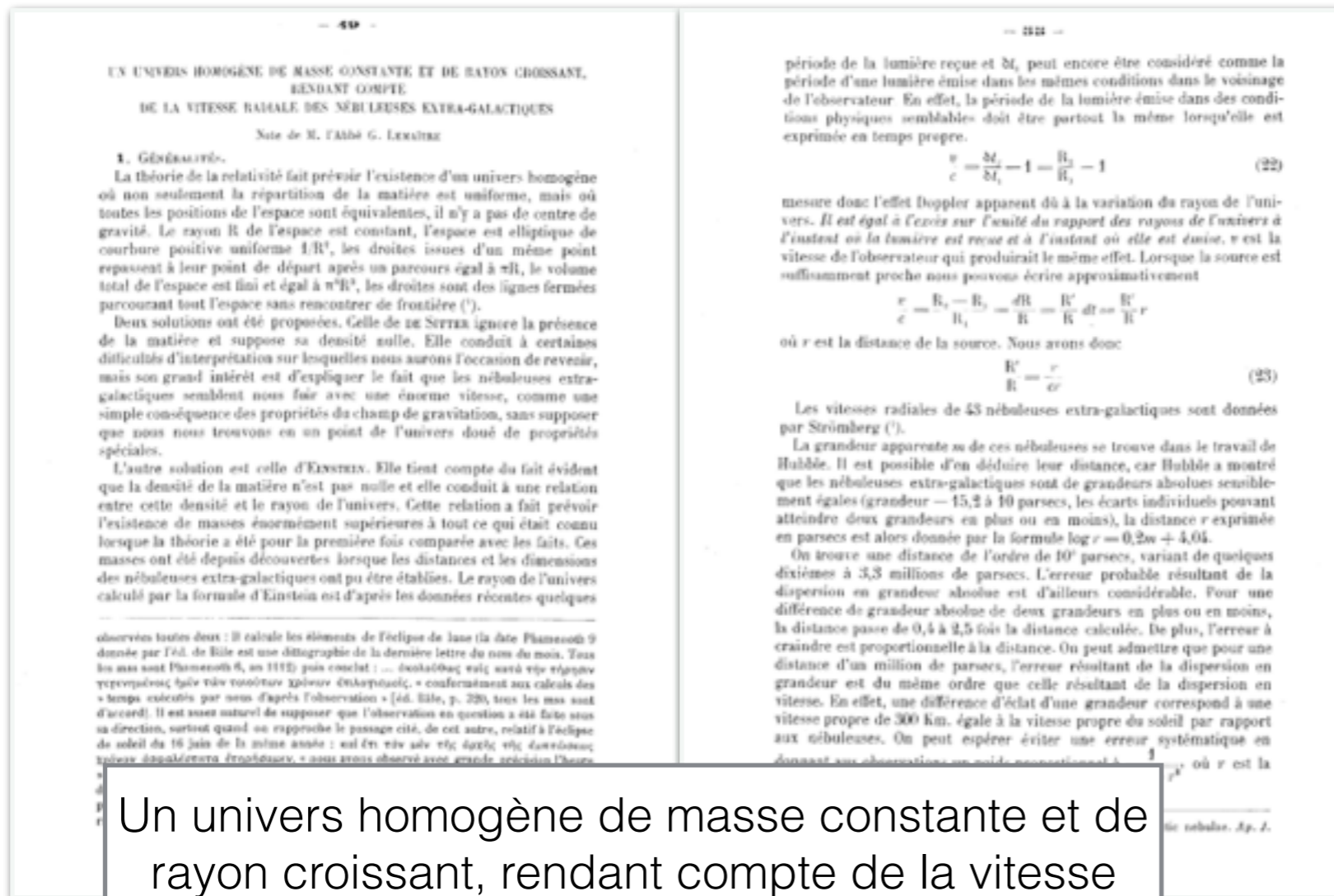


<https://www.museumofflight.org/iya-hubble>



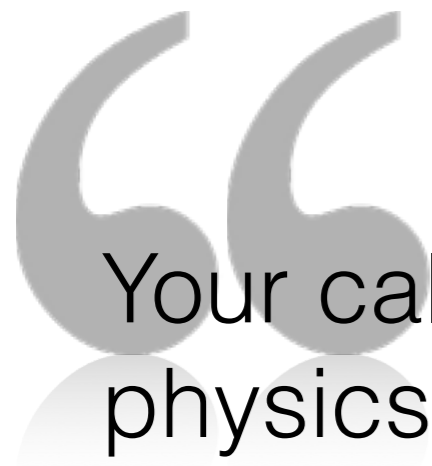
# Lemaître's wet (1927)

De radiale snelheid van de nebulae neemt toe met de afstand



Un univers homogène de masse constante et de rayon croissant, rendant compte de la vitesse radiale des nébuleuses extra-galactiques.  
*Annales de la Société scientifique de Bruxelles*





Your calculations are correct, but your grasp of physics is abominable. *Einstein, 1927.*



<http://www.astrosurf.com/luxorion/Illustrations/einstein-lemaître-pasadena-1932.jpg>

# Hubbles publicatie (1929)

168

ASTRONOMY: E. HUBBLE

Proc. N. A. S.

appearance the spectrum is very much like spectra of the Milky Way clouds in Sagittarius and Cygnus, and is also similar to spectra of binary stars of the W Ursae Majoris type, where the widening and depth of the lines are affected by the rapid rotation of the stars involved.

The wide shallow absorption lines observed in the spectrum of N. G. C. 7619 have been noticed in the spectra of other extra-galactic nebulae, and may be due to a dispersion in velocity and a blending of the spectral types of the many stars which presumably exist in the central parts of these nebulae. The lack of depth in the absorption lines seems to be more pronounced among the smaller and fainter nebulae, and in N. G. C. 7619 the absorption is very weak.

It is hoped that velocities of more of these interesting objects will soon be available.

## A RELATION BETWEEN DISTANCE AND RADIAL VELOCITY AMONG EXTRA-GALACTIC NEBULAE

By EDWIN HUBBLE

MOUNT WILSON OBSERVATORY, CARNEGIE INSTITUTION OF WASHINGTON

Communicated January 17, 1929

Determinations of the motion of the sun with respect to the extra-galactic nebulae have involved a  $K$  term of several hundred kilometers which appears to be variable. Explanations of this paradox have been sought in a correlation between apparent radial velocities and distances, but so far the results have not been convincing. The present paper is a re-examination of the question, based on only those nebular distances which are believed to be fairly reliable.

Distances of extra-galactic nebulae depend ultimately upon the application of absolute-luminosity criteria to involved stars whose types can be recognized. These include, among others, Cepheid variables, novae, and blue stars involved in emission nebulosity. Numerical values depend upon the zero point of the period-luminosity relation among Cepheids, the other criteria merely check the order of the distances. This method is restricted to the few nebulae which are well resolved by existing instruments. A study of these nebulae, together with those in which any stars at all can be recognized, indicates the probability of an approximately uniform upper limit to the absolute luminosity of stars, in the late-type spirals and irregular nebulae at least, of the order of  $M$  (photographic) =  $-6.3$ .<sup>1</sup> The apparent luminosities of the brightest stars in such nebulae are thus criteria which, although rough and to be applied with caution,



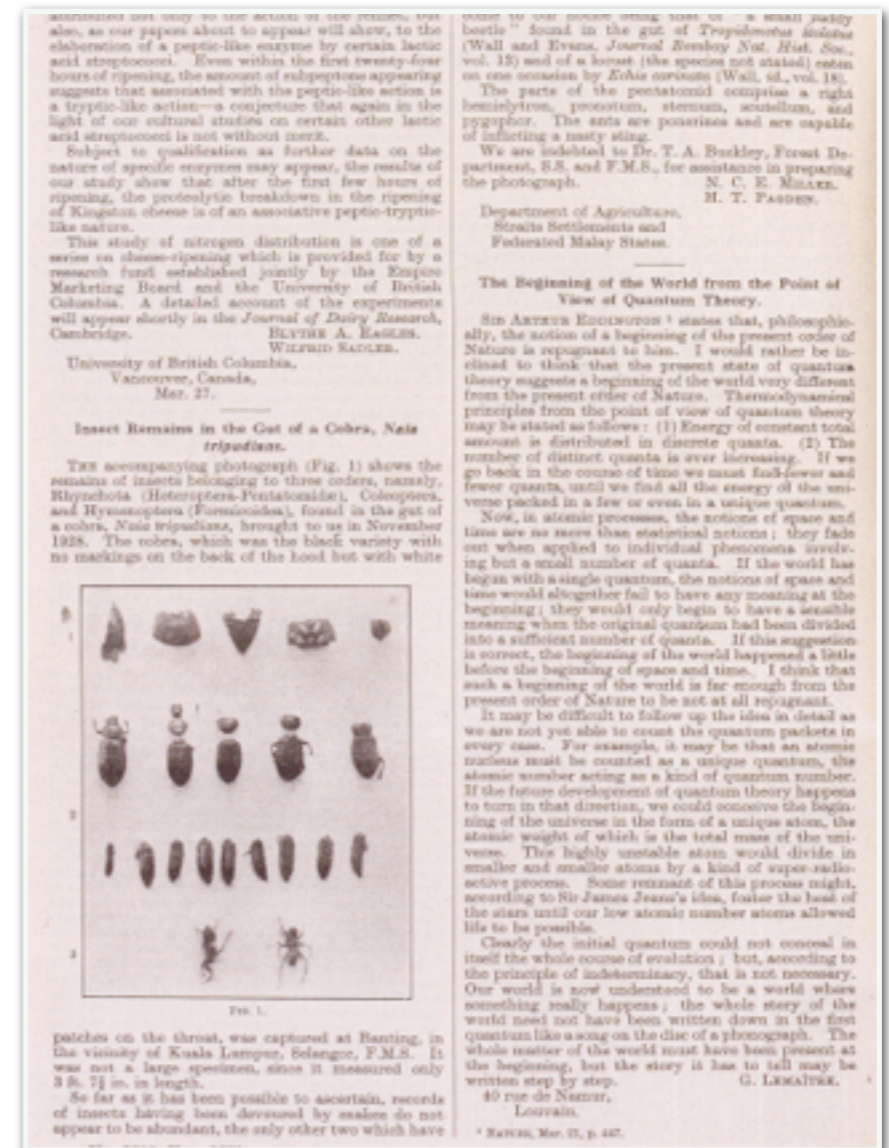
<http://www.interactions.org/quantumuniverse/qu2006/discovering/einstein.html>



# The beginning of the world from the point of view of quantum theory

## Lemaître's oerknalmodel

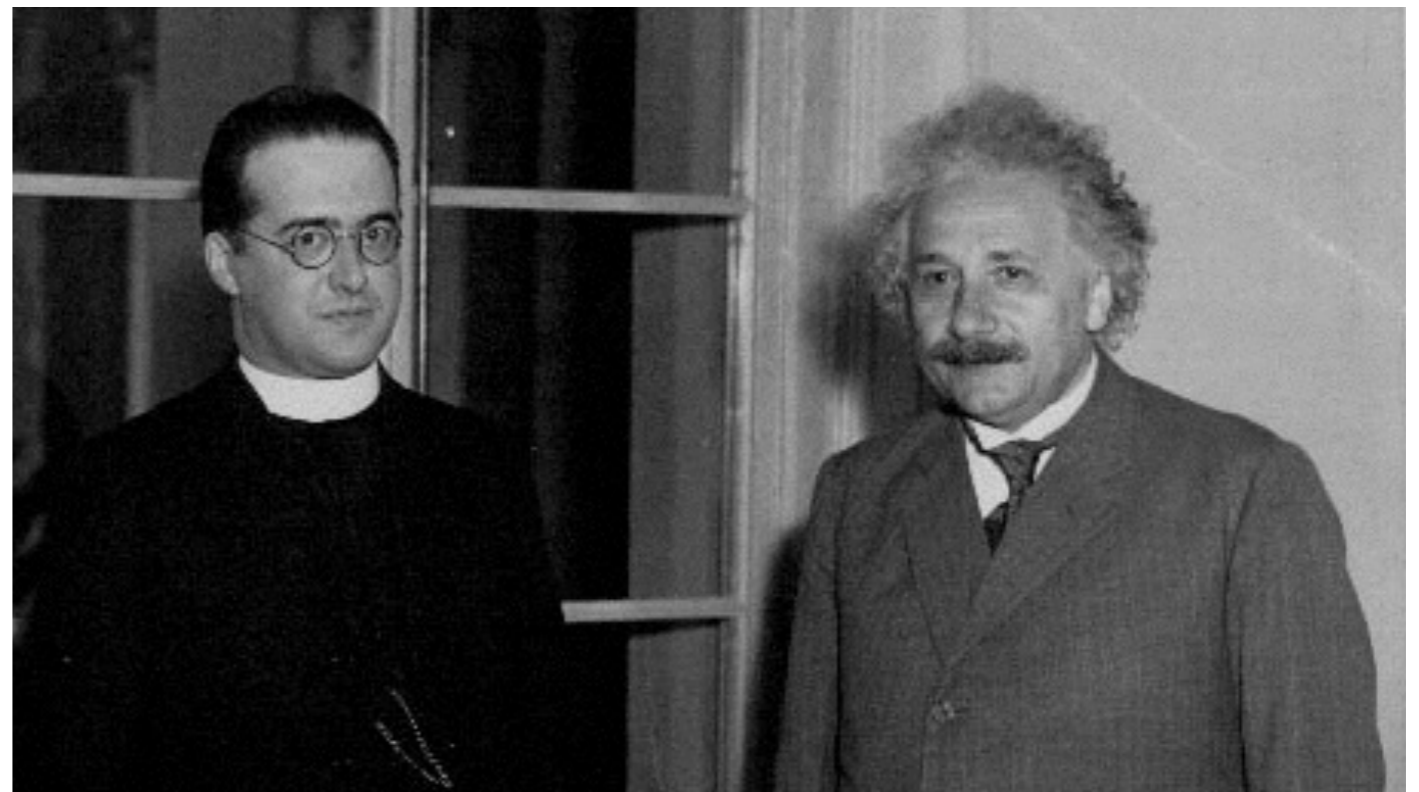
If the world has begun with a single quantum, the notions of space and time would altogether fail to have a meaning at the beginning. [...] *If this suggestion is correct, the beginning of the world happened a little before the beginning of space and time.* I think that such a beginning of the world is far enough from the present order of Nature to be not at all repugnant.



Nature 127, 706 (9 May 1931)

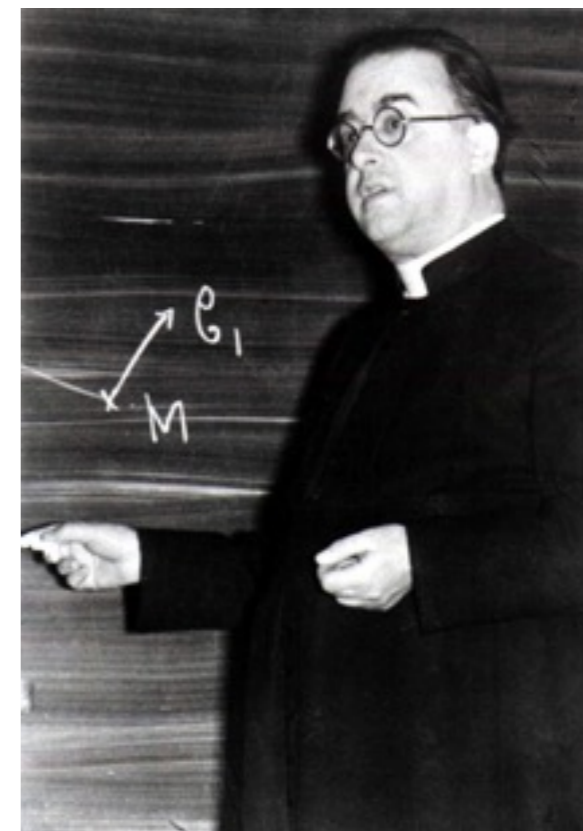
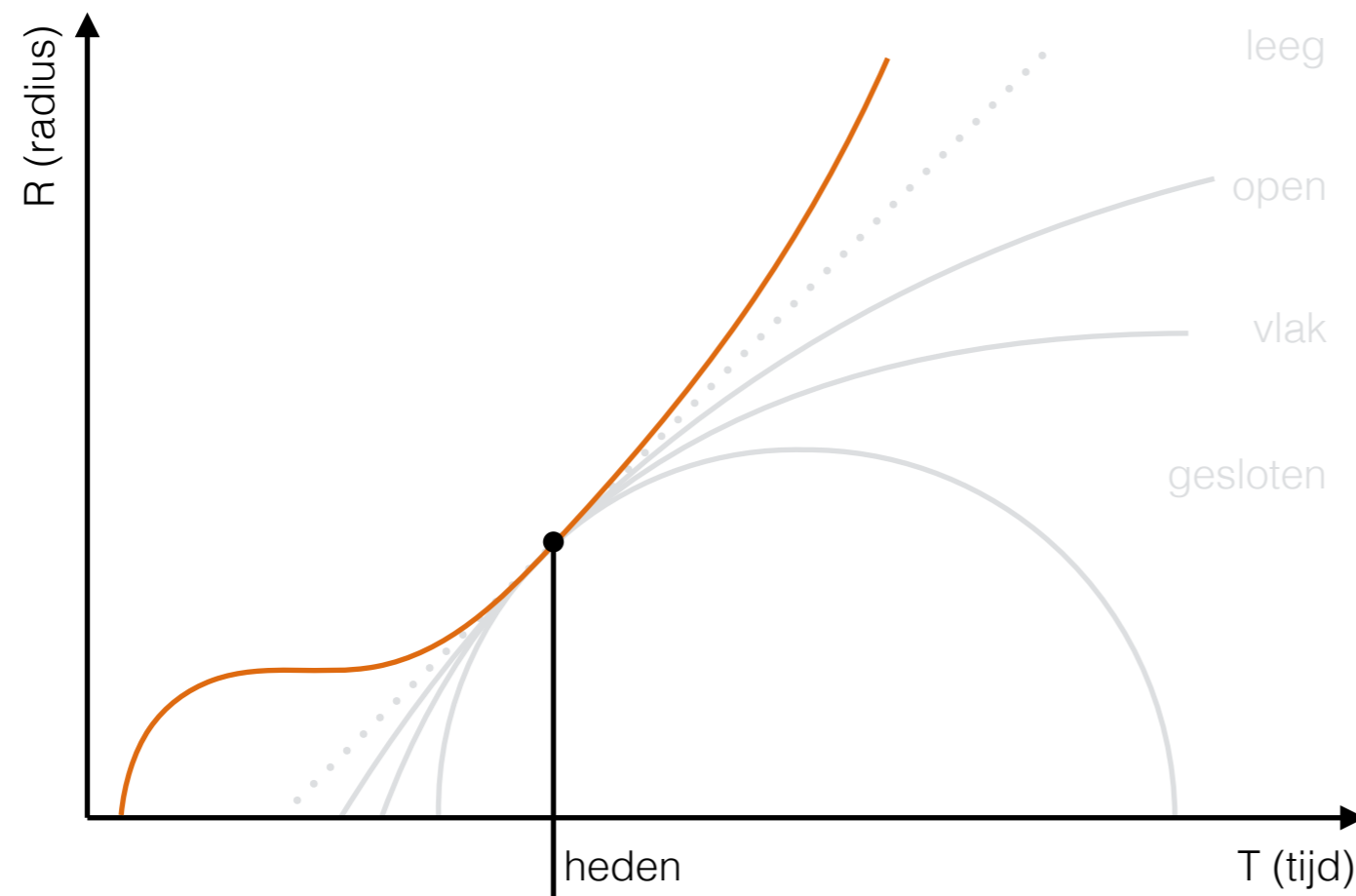


This is the most beautiful and satisfactory explanation of creation to which I have ever listened. *Einstein, 1933*



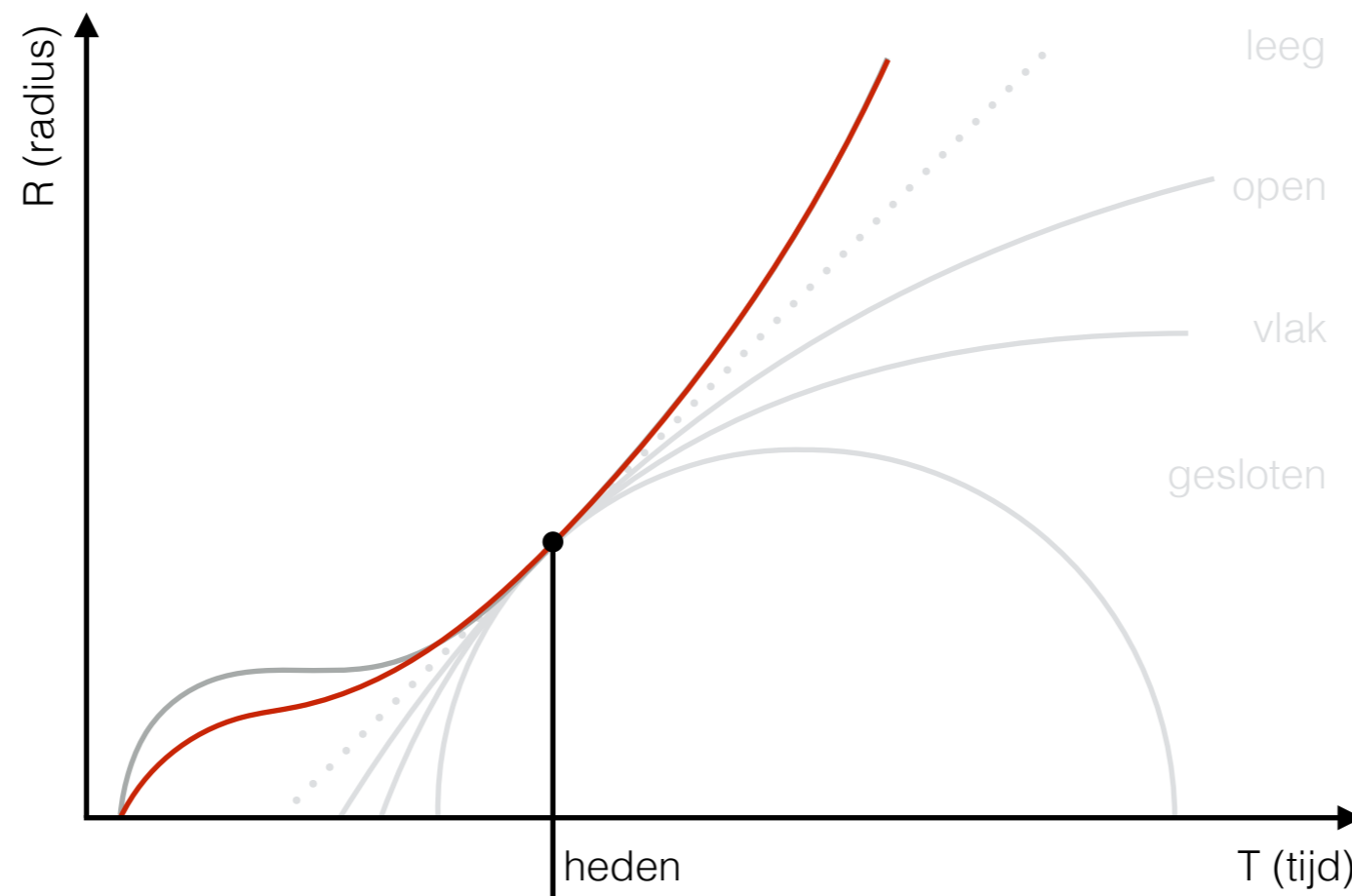
[http://www.astr.ucl.ac.be/images/abouttheinstitute/Lemaitre\\_Einstein.jpg](http://www.astr.ucl.ac.be/images/abouttheinstitute/Lemaitre_Einstein.jpg)

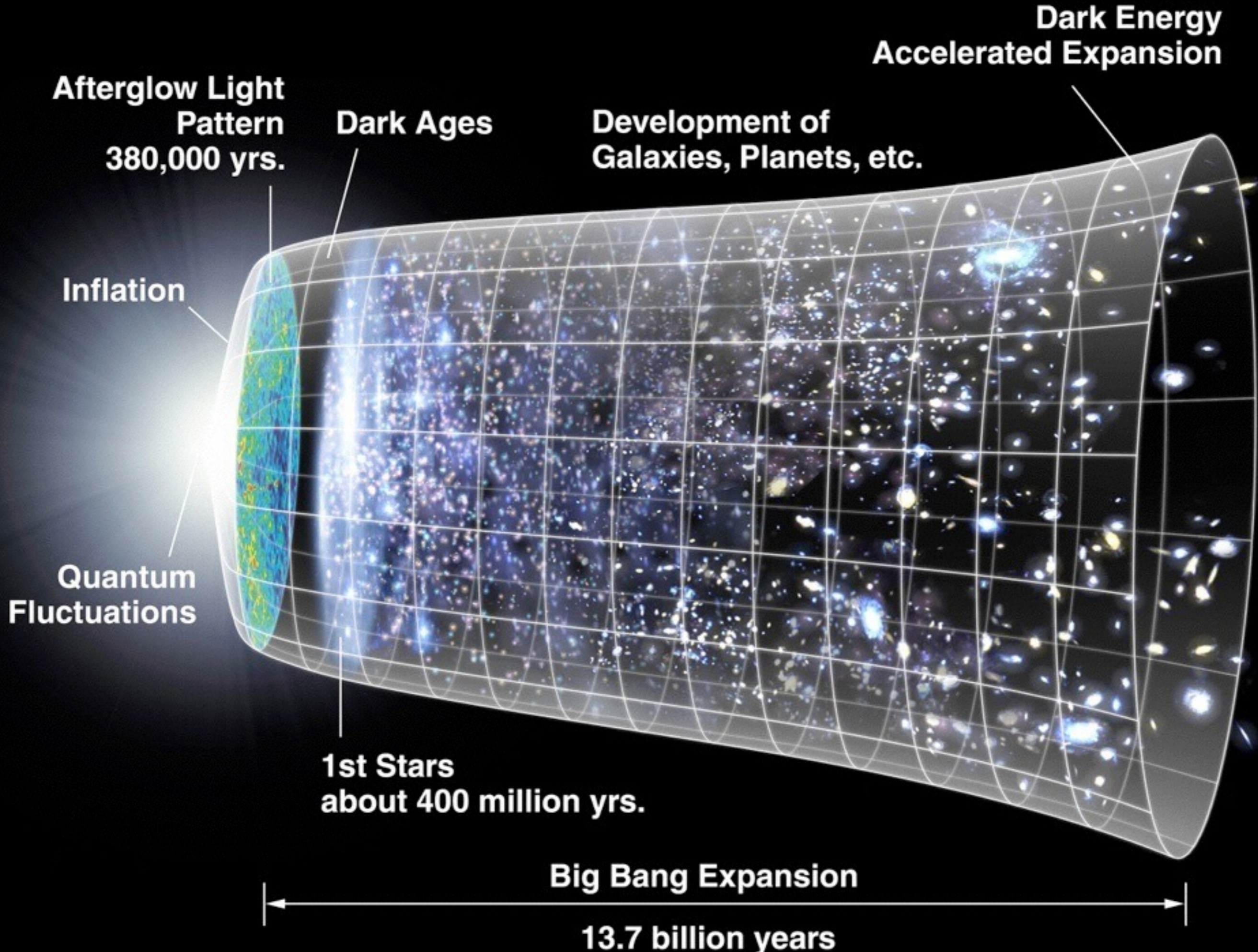
# Model van Lemaître



<http://www.zaujimavysvet.sk>

# Concordantiemodel

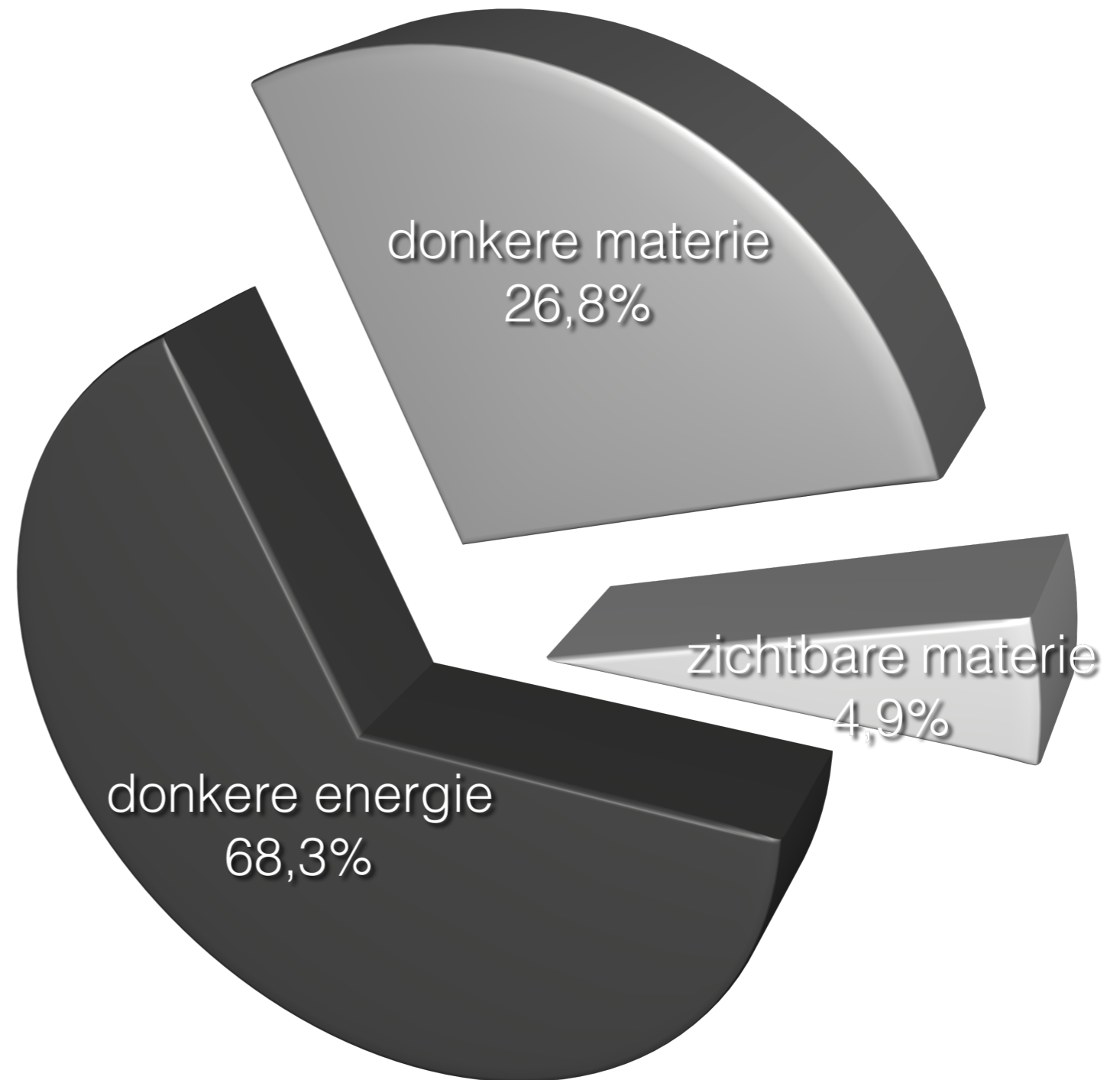




# $\Lambda$ CDM

## *Lambda Koude Donkere Materie*

Lambda speelt een rol:  
vertegenwoordigt de  
invloed van donkere  
energie



# $\Lambda$ CDM

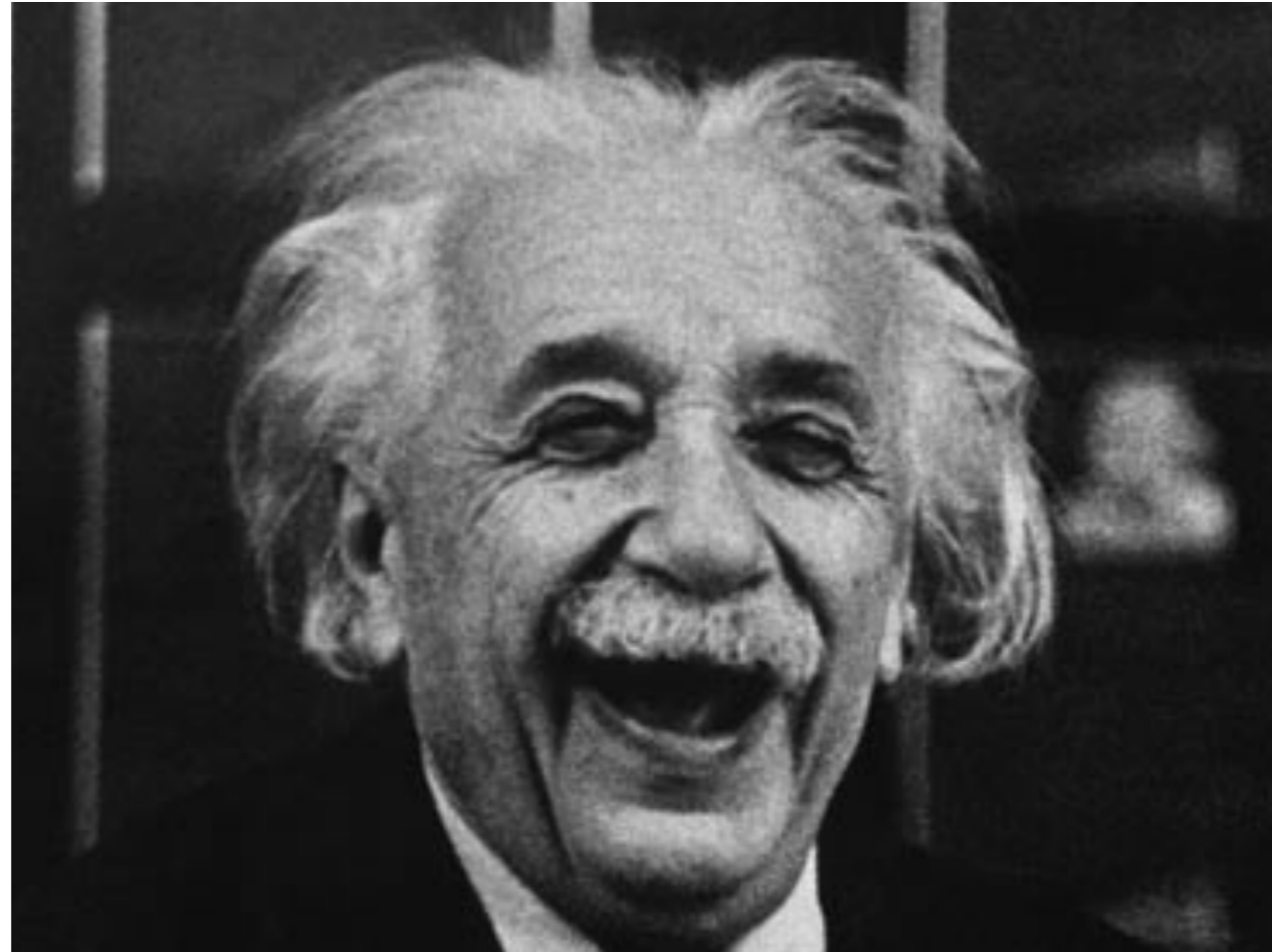
- terugkijktijd  $\sim 13,750 \times 10^9$  jaar
- samenstelling
  - $\sim 5\%$  baryonische materie (atomen)
  - $\sim 27\%$  donkere materie
  - $\sim 68\%$  donkere energie (aanvankelijk)
- kosmologische parameters drukken de karakteristieken van het heelal uit:
  - $\Omega_\Lambda = 0,7$  – kan positief of negatief zijn (antigravitatie of extra gravitatie)
  - $\Omega_M = 0,3$  – bijdrage van de massa (baryonische en donkere materie)
  - $H_0 = \sim 70$  km/s/Mpc (constante van Hubble)
  - $q < 0$  (vertraging of versnelling)

# Einsteins blunder?

Einstein zou volgens George Gamow (1956) de invoering van  $\lambda$  de grootste blunder van zijn leven genoemd hebben. Er is geen andere bron.  
(Mario Livio)


Einstein quickly recognized the importance of this discovery. In the last edition of his book *The Meaning of Relativity* he wrote: "The mathematician Friedman found a way out of this dilemma. He showed that it is possible, according to the field equations, to have a finite density in the whole (three-dimensional) space, without enlarging these field equations ad hoc." Einstein remarked to me many years ago that the cosmic repulsion idea was the biggest blunder he had made in his entire life.



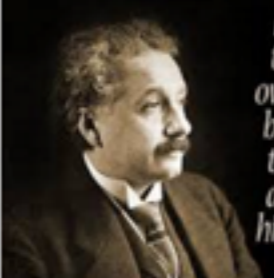


De geciteerde Einstein

"Everybody is a genius. But if you judge a fish by its ability to climb a tree, it will live its whole life believing that it is stupid."  
- Albert Einstein




Everyone sits in the prison of his own ideas; he must burst it open, and that in his youth, and so try to test his ideas on reality.  
-ALBERT EINSTEIN



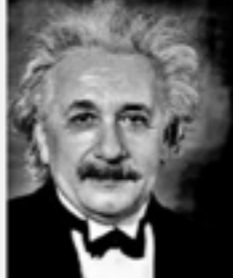
"I NEVER SAID HALF THE CRAP PEOPLE SAID I DID"  
ALBERT EINSTEIN



Example ISN'T ANOTHER WAY TO TEACH, IT IS THE ONLY WAY TO TEACH



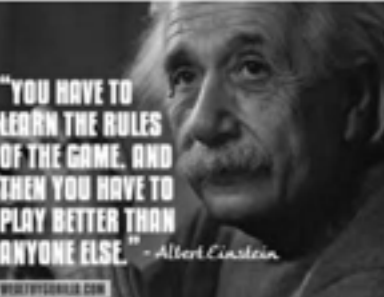
"The world is a dangerous place to live; not because of the people who are evil, but because of the people who don't do anything about it."  
Albert Einstein



THE TRUE SIGN OF INTELLIGENCE IS NOT KNOWLEDGE BUT IMAGINATION.  
Albert Einstein  
German Theoretical Physicist (1879-1955)




"YOU HAVE TO LEARN THE RULES OF THE GAME, AND THEN YOU HAVE TO PLAY BETTER THAN ANYONE ELSE."  
- Albert Einstein



A PERSON WHO NEVER MADE A MISTAKE NEVER TRIED ANYTHING NEW.  
ALBERT EINSTEIN




"I have no special talents. I am only passionately curious."  
- Albert Einstein




"A person who is part of a world, called by us the 'universe,' a part limited in time and space, his consciousness expands, his thoughts and feelings, as if he were not limited by the boundaries of the universe. This expansion is a kind of freedom for us, because it is not limited by space and time, and it is not limited by the boundaries of the universe. It is this freedom that makes it possible for us to think about the universe as a whole, and to see it as a part of a larger whole."  
- Albert Einstein



To punish me for my contempt of authority, Fate has made me an authority myself.  
-ALBERT EINSTEIN




"The world is not dangerous because of those who do harm but because of those who look at it without doing anything."  
- Albert Einstein



For what he discovered a mass of ignorance had hidden in himself a mass of wisdom.



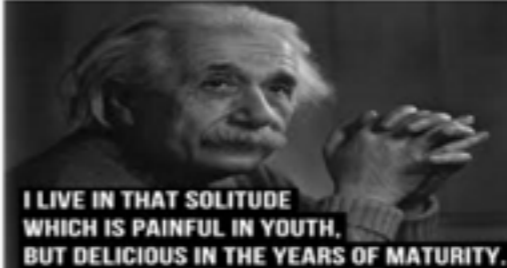
"You have to learn the rules of the game. And then, you have to play better than anyone else."  
-Albert Einstein




Learn from yesterday, live for today, hope for tomorrow. The important thing is not to stop questioning.  
Albert Einstein




I LIVE IN THAT SOLITUDE WHICH IS PAINFUL IN YOUTH, BUT DELICIOUS IN THE YEARS OF MATURITY.  
ALBERT EINSTEIN




KNOWLEDGE IS LIMITED, IMAGINATION ENRICHES THE WORLD.



"NEVER LOSE A HOLY CURIOSITY."  
ALBERT EINSTEIN



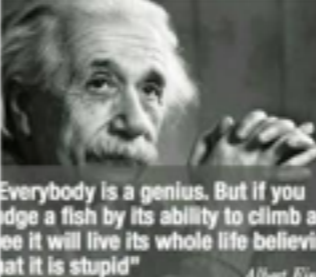
I fear the day when people post me on the internet with I've never said.  
-Albert Einstein




De mens die zich omringt met feiten, niet toestaat dat hij verrast wordt, geen flits van intuïtie kent, geen veronderstellingen maakt, geen risico neemt, leeft in een afgesloten cel.  
Albert Einstein



Everybody is a genius. But if you judge a fish by its ability to climb a tree it will live its whole life believing that it is stupid"  
Albert Einstein



If you can't explain it simply, you don't understand it well enough.  
- Albert Einstein



GREAT SPIRITS HAVE ALWAYS ENCOUNTERED VIOLENT OPPOSITION FROM MEDIOCRE MINDS  
ALBERT EINSTEIN






... da könnt' mir halt der liebe Gott leid tun,  
die Theorie stimmt doch.

–Einstein, controversieel citaat, via Ilse Rosenthal-Schneider



I want to know how God created this world. I'm not interested in this or that phenomenon, in the spectrum of this or that element. I want to know His thoughts; the rest are details.

–Einstein, controversieel citaat, via Esther Salaman



While it is true that scientific results are entirely independent from religious or moral considerations, those individuals to whom we owe the great creative achievements of science were all of them imbued with the truly religious conviction that this universe of ours is something perfect and susceptible to the rational striving for knowledge.

Einstein, A., 1941, "Science, Philosophy and Religion, A Symposium", *Conference on Science, Philosophy and Religion in Their Relation to the Democratic Way of Life*, New York.

# Nog eens alles lezen?



